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27                   **Response of Fogo volcano (Cape Verde) to lunisolar gravitational forces**  
28   **during the 2014-2015 eruption**

29  
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41  
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43   **Abstract**

44   Volcanoes are complex systems that evolve in space and time as a result of their eruptive activity.

45   Volcanic eruptions represent the ultimate expression of a complex interplay between internal and

46   external processes that span across different time scales. Deciphering how internal and external

47   processes interact at the time scale of eruptions may provide key insights on the temporal evolution

48   of eruptions and also help to better evaluate associated volcanic hazards. Studies of the tidal

49   influence on volcanic activity have fallen within this context, although the cause-effect relationship

50   between tides and eruptions is still unclear. In this study, we used Singular Spectrum Analysis to

51   analyze three time-series, namely the seismic tremor, SO<sub>2</sub> emission and lava volume flow rate,

52   which cover the first month of effusive activity at Fogo volcano, Cape Verde, in 2014-2015. We

53   detect 9 tidal periodicities and up to 5 in each time-series ranging from semi-diurnal to fortnightly

54 periods. We show that the movement of magma at crustal depths and at surface as well as gas  
55 emission during the effusive eruption are all modulated by lunisolar gravitational forces. We  
56 highlight the relevance of the volcano location on Earth, which together with the timing of the  
57 eruption, associated with a specific astronomical configuration, result in a specific combination of  
58 tides that directly influence the volcano eruptive activity. With this data set, we further investigate  
59 the response of Fogo volcano to this external forcing. We show that during the 2014-2015 eruption,  
60 Fogo volcano acted as a bandpass filter to quasi-permanent tidal oscillations.

61

62 **Keywords:** 2014-2015 Fogo eruption; earth tides; SSA; seismic tremor; eruptive activity; external  
63 forcing

64

## 65 **1. Introduction**

66 Periodicities corresponding to lunisolar tides have for long been observed in seismic and volcanic  
67 activity (Mauk and Johnston, 1973; Heaton, 1975; Sparks, 1981; Rymer and Brown, 1984; Bodri  
68 and Iizuka, 1989; Ciglioni et al., 2009; Bredemeyer and Hansteen, 2014; Delorey et al., 2017;  
69 Petrosino et al., 2018; Ricco et al., 2019; Varga and Grafarend, 2019). In addition, recent studies  
70 have evidenced presence of tides at time scales ranging from a few days to decades, in the polar  
71 motion (Lopes et al., 2017), plate motion (Zaccagnino et al., 2020) and even in climate indexes as a  
72 response to the temporal evolution of atmospheric pressures (Le Mouél et al., 2019a). At volcanoes,  
73 periodic variations have been revealed by various kind of observations such as lava lake height,  
74 degassing, micro-seismicity and Long-Period (LP) events, volcanic tremor, ground tilt, temperature  
75 of fumarolic fields, energy radiated by lava or strength of eruptive phases (Golombek and Carr,  
76 1978; Berrino et al., 1991; Williams-Jones et al., 2001; Custodio et al., 2003; Sottili et al., 2007;  
77 Ciglioni et al., 2009; De Lauro et al., 2012, 2013, 2018; Sottilli and Palladino, 2012; Bredemeyer  
78 and Hansteen, 2014; Girona et al., 2018; Dinger et al., 2018; Caputo et al., 2020; Dumont et al.,  
79 2020; Petrosino et al., 2020). Lunisolar gravitational forces have also been evoked as triggers of

80 volcanic eruptions (Mauk and Johnston, 1973; Dzurizin, 1980; Jentzsch et al. 2001; Dumont et al.,  
81 2020).

82 Different approaches have been considered to demonstrate the correlation between tidal  
83 action and variations in volcanic activity, from statistics to spectral or principal component analysis  
84 (see examples in Mauk and Johnston, 1973; Patanè et al., 1994; Girona et al., 2018; De Lauro et al.,  
85 2013; Dumont et al., 2020). The main and most frequent periodicities detected in time-series  
86 acquired at volcanoes are those of the fortnightly, which is induced by the Moon's declination, the  
87 semi-diurnal caused by the Sun and lunar elliptic trajectory, as well as the diurnal component which  
88 is related to the Sun (Berrino et al., 1991; Custodio et al., 2003; Sottilli and Palladino, 2012; De  
89 Lauro et al., 2018; Girona et al., 2018; Le Mouél et al., 2019b). These tidal oscillations are,  
90 however, not the only to modulate activity at volcanoes, as suggested by a couple of studies (Sottilli  
91 and Palladino, 2012; Bredemeyer and Hansteen, 2014; Caputo et al., 2020; Dumont et al., 2020). In  
92 spite of all these studies, providing clear evidence on the correlation between processes that occur at  
93 very different time scales has proved to be particularly challenging. In addition and beyond the  
94 techniques used, the cause-effect relationship between tidal action and volcanic activity has  
95 remained elusive (Sparks, 1981; Neuberg, 2000). Tidal stresses and tidal acceleration have been  
96 suggested as the main drivers of tidal forcing although tidal stresses are about 3-5 orders of  
97 magnitude smaller than tectonic stresses (Mauk and Johnston, 1973; Sparks, 1981), which makes  
98 them too low to induce fracturing (McMillan et al., 2019; Dumont et al., 2020). The complexity of  
99 volcanic systems, which is mainly due to their eruptive history, composition, internal structure and  
100 tectonic setting, also plays an important role on how volcanoes respond to the quasi-permanent tidal  
101 oscillations explaining also why all volcanoes do not show a similar sensitivity to Earth tides (Mauk  
102 and Johnston, 1973; Dzurizin, 1980). Additionally, the specific positions of volcanoes on the planet,  
103 means they are not influenced similarly by the different Earth tides, as the zonal, tesseral and  
104 sectorial components of the tidal potential have a specific spatial distribution on Earth (see Jobert  
105 and Coulomb, 1973; chapter 18). The spatial distribution and especially the latitudinal position of

106 the different gravitational bodies that interact with the volcano is determinant in the way the tidal  
107 potential applies on Earth (Mauk and Johnston, 1973; Hamilton, 1973; Jobert and Coulomb, 1973) .

108 In this study, we adopted a similar approach to Dumont et al., (2020) and used the Singular  
109 Spectrum Analysis to investigate three geophysical time-series, namely the SO<sub>2</sub> emission, lava  
110 volume flow rate (VFR) and seismic tremor, spanning the 2014-2015 mix eruption at Fogo volcano,  
111 Cape Verde. This eruption started on November 23<sup>rd</sup>, 2014 and ended on February 8<sup>th</sup>, after 2.5  
112 months of intense effusive activity (Mata et al., 2017; Richter et al., 2016). The resulting extensive  
113 lava field ( $43.7 \times \pm 5.2 \cdot 10^6 \text{ m}^3$ ) covered 4.84 km<sup>2</sup> of Chã das Caldeiras and destroyed two villages  
114 (Richter et al., 2016; Jenkins et al., 2017). We characterize the lunisolar tidal influence on this  
115 dominantly effusive eruption which took place in the Equatorial zone, to further investigate the  
116 specific response of Fogo volcano to this permanent external forcing associated with Earth tides.

117

## 118 **2. Fogo volcano and the 2014-2015 eruption**

119 Fogo volcano is located on one of the dozen islands composing the Cape Verde archipelago,  
120 which lies ~450 km west of Africa, in the North Atlantic (Figure 1). The west-open crescent formed  
121 by the islands rises on the top of a 1200 km diameter swell characterized by the largest bathymetric  
122 (+ 2 km) and geoid (+ 8m) anomalies in the world which are thought to result from an upwelling  
123 mantle plume (Courtney and White, 1986; Grevemeyer, 1999; Carvalho et al., 2019). The rise of the  
124 Cape Verde islands initiated early Miocene (~22 Ma) but most of the islands formed in the last 16  
125 Ma (Holm et al., 2008). Mantle channeling has been proposed to explain the position of islands  
126 along three main groups: northwest, east and southwest (Holm et al., 2008). Fogo belongs to the  
127 latest group and is one of the most recent islands formed in the Quaternary (Ramalho et al., 2010).  
128 Fogo and its western neighboring island, Brava, are the center of the present-day seismic and  
129 magmatic activity of the whole archipelago (Silva et al., 1999; Amelung and Day, 2002; Custodio et  
130 al., 2003; Dionis et al., 2015; Eisele et al., 2015; González et al., 2015; Faria and Fonseca, 2014;  
131 Vales et al., 2014; Leva et al., 2019).

132 Fogo island is characterized by steep slopes topped by a 9 km diameter horseshoe-shaped  
133 caldera, Chã das Caldeiras. The geometry of the 1-km high caldera wall evidences past collapse  
134 events whose nature and origin are still debated nowadays although the risk of associated tsunami is  
135 well recognized. One catastrophic event was identified at ~65 to 84 kyrs but it may not be the only  
136 one (Day et al., 1999; Ramalho et al., 2015; Barrett et al., 2019). Since the last flank collapse, the  
137 volcanic activity has been focused on the Fogo volcano, after which the island was renamed, a  
138 stratovolcano rising at 2829 m above sea level nearby the center of the caldera (Figure 1). The Fogo  
139 volcano is considered as one of the most active volcanoes in the Atlantic, with 28 eruptions since  
140 the settlement in the early 15th century (Faria, 2010). The repose period between eruptions has  
141 changed over the last 600 years with eruptions every ~60 years on average between 1400 and 1680,  
142 one every ~5 years between 1750 and 1875 and one every ~20 years since the last century (Faria,  
143 2010). Tephrostratigraphic analyses of off-shore tephra revealed a past and highly explosive  
144 eruptive activity of Fogo estimated to occur every ~3000 years during the last 150 kyrs (Eisele et  
145 al., 2015). The last summit eruption occurred in 1785 and since then most eruptions took place on  
146 volcano flanks (Faria, 2010; Amelung and Day, 2002; González et al., 2015). The last three  
147 eruptions (1951, 1995, 2014-15) produced extensive lava flows that threatened local populations  
148 and caused the total destruction of two villages in 2014-15 (Richter et al., 2016; Jenkins et al.,  
149 2017).

150 The last eruption of Fogo started on 23 November 2014 and lasted 77 days, ending on 7  
151 February 2015. The eruptive fissure opened with a violent explosion a few hundred meters away  
152 from the 1995 cinder cone. It was initially formed by 6-7 individual vents that reduced to 3 over the  
153 course of the eruption. Lava fountains and Strombolian explosions mainly composed this volcanic  
154 activity punctuated by stronger explosions and ash plume emissions (Mata et al., 2017; Calvari et  
155 al., 2018). Remote-sensing imagery revealed a mean effusion rate of  $6.8 \text{ m}^3 \cdot \text{s}^{-1}$ , with a peak of ~24-  
156  $27 \text{ m}^3 \cdot \text{s}^{-1}$  achieved on 24 November and very low discharge rates after late December (Cappello et  
157 al., 2016; Calvari et al., 2018). The new lava field covered an area of  $\sim 5.42 \text{ km}^2$  at the end of the

158 eruption, corresponding to a total volume of  $9.6\text{-}18.2 \cdot 10^6 \text{ m}^3$  (Bagnardi et al., 2016; Richter et al.,  
159 2016). The magma erupted during this eruption was mainly alkaline with some evolved  
160 compositions, e.g. tephrites to phonotephrites. The Nb/U ratios, lower than those of the two  
161 previous eruptions, were interpreted as an expression of small-scale mantle heterogeneity (Mata et  
162 al., 2017).

163

### 164 **3. Data and Methods**

#### 165 *3.1 Retrieval of lava and SO<sub>2</sub> emissions*

166 Lava Volume Flow Rate (VFR) and SO<sub>2</sub> vertical column densities (VCD) have been retrieved  
167 together from the HOTVOLC system. HOTVOLC is a Web-GIS volcano monitoring system using  
168 SEVIRI (Spinning Enhanced Visible and Infrared Imager) sensor on-board Meteosat geostationary  
169 satellite (<https://hotvolc.opgc.fr>) and developed at the OPGC (Observatoire de Physique du Globe  
170 de Clermont-Ferrand) in 2009 (Gouhier et al., 2016). The spectral bands of the SEVIRI sensor  
171 allow us to simultaneously characterize volcanic ash (Guéhenneux et al., 2015), sulfur dioxide  
172 (Gauthier et al., 2016; Gouhier and Paris, 2019), and lava flow emissions (Gouhier et al., 2012;  
173 Gouhier et al., 2016; Thivet et al., 2020). It is designed for the real-time monitoring of ~50 active  
174 volcanoes, including Fogo volcano and provides high value-added products at the frequency of 1  
175 image / 15 min with a pixel resolution of  $3 \text{ km} \times 3 \text{ km}$  at nadir. For this study, 7488 multispectral  
176 IR images have been processed, covering the whole eruption duration.

177

178 Lava effusion rates have been estimated using instantaneous lava volumes, calculated from infrared  
179 techniques. This method lies on the mass conservation principle and benefits from a rapid succes-  
180 sion of images obtained using geostationary satellites. The physical principle is as follows: the  
181 pixel-integrated thermal anomaly measured at a given instant is the balance of contributions related  
182 to (i) hot lava material newly emplaced (at time  $t$ ) and (ii) cooling lava material previously em-

183 placed (at time  $t-I$ ). This problem can be addressed using the mass conservation principle following  
184 the differential equation:

$$\frac{dV(t)}{dt} = Q(t) - kV(t)$$

185  
186  
187  $V$  is the total volume of lava in a given image,  $Q$  is the lava volume flow rate (i.e., source term) and  
188  $-kV$  is the loss term with  $k$  representing the lava cooling rate ( $s^{-1}$ ). Note that satellite observations  
189 provide discrete time series, to obtain  $Q$  we thus solve analytically the above differential equation  
190 with  $\Delta t$  being the time interval between two consecutive images such that:

$$Q(t) = k \frac{V_i - V_{i-1} e^{-k\Delta t}}{1 - e^{-k\Delta t}}$$

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192  
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194  
195 The cooling rate  $k$  is often expressed as  $k=I/\tau$ , where  $\tau$  is the lava e-cooling time. This method,  
196 referred to as the Volume Flow Rate (VFR) method, allows the calculation of quasi-instantaneous  
197 lava effusion rates (in  $m^3/s$ ). The quantification of instantaneous lava effusion rates is essential as it  
198 strongly controls the lava front speed and flow area. Moreover, the short-term evolution of lava  
199 effusion rate traduces eruptive dynamics changes from shallow depth to the surface.

200  
201 Volcanic  $SO_2$  emissions have long been characterized from thermal infrared sensors using  
202 absorption features of either the 8.7- $\mu m$  waveband (e.g., [Watson et al., 2004](#)) or the 7.3- $\mu m$   
203 waveband ([Prata et al., 2003](#)). The retrieval scheme used here follows from [Realmuto et al., \(1994\)](#),  
204 where the  $SO_2$  column abundance, referred to as VCD, is estimated from the 8.7- $\mu m$  waveband us-  
205 ing a linear regression derived from a least square fitting procedure between sensor measurements  
206 (i.e., SEVIRI) and simulated radiances using the Modtran radiative transfer model. VCD are usually  
207 expressed in Dobson Unit (DU) for  $SO_2$  abundance. One Dobson Unit corresponds to  $2.69 \times 10^{20}$   
208 molecules per square meter, i.e.,  $2.86 \times 10^{-5} \text{ kg/m}^2$ . Finally, we provide here the maximum  $SO_2$  VCD  
209 (in DU) for each image in the form of a time series over the whole eruption duration.

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### 213 *3.2 Seismic tremor*

214 The amplitude of seismic tremor was quantified using data collected by a temporary network of  
215 seven broadband seismic stations (Figure 1), which were quickly deployed after the eruption began.  
216 This network operated for about one month and a half, between November 28, 2014 and January  
217 15, 2015. All stations were equipped with CMG6TD sensors (30s to 100 Hz) and recorded  
218 continuously at a sample rate of 100 Hz.

219 Continuous seismic data were initially instrument-corrected and converted to true ground velocity,  
220 after demeaning, detrending and tapering. Data were then filtered between 1 and 4 Hz, the  
221 frequency band typically associated to volcanic tremor. In order to assess the evolution of the  
222 amplitudes of seismic tremor, we computed the Root Median Square (RMeS) of the band-pass  
223 filtered continuous seismic records as in Eibl et al., (2017). RMeS is calculated as  
224  $\sqrt{\text{Median}(y_i^2, \dots, y_j^2)}$ , where the median is calculated based on the squared amplitudes of seismic  
225 data recorded in each 15-min sliding window ( $y_i$  and  $y_j$  are the initial and final seismic amplitudes  
226 in each window). We use a 50% overlap between neighbor windows. For this study, we considered  
227 the seismic tremor observed at station FCVT (Covatina), which is the station closest to the eruptive  
228 fissure (< 5 km, Figure 1).

229

### 230 *3.3 Length-of-Day (l.o.d.) and sea level measurements*

231 The length-of-day (l.o.d.) is an astronomical measure of the Earth's rotation velocity. It is one of the  
232 Earth orientation parameters, measured from the combination of Very Long Baseline Radio  
233 Interferometry (VLBI), Satellite Laser ranging (LSR) with Doppler Orbitography and Radio  
234 positioning Integrated by Satellite (DORIS) techniques (Bizouard et al., 2019). For this study, we  
235 used the EOP14C04 data set provided by the International Earth Rotation Service

236 (<https://www.iers.org/IERS/EN/DataProducts/EarthOrientationData/eop.html>, IERS, Paris, France). Both regular  
237 and irregular l.o.d. variations have been evidenced at several time scales, from a few days to  
238 decades (Guinot, 1973; Lambeck et al., 2005; Ray and Erofeeva, 2014; Le Mouél et al., 2019b).  
239 Decadal fluctuations have been attributed to core/mantle interaction and their dynamics, while rapid  
240 variations to external forcings, including atmosphere dynamics and Earth tides (Lambeck et al.,  
241 2005). On the time scale of a few months, l.o.d. variations do not exceed 2 ms.  
242 We complement l.o.d. daily measurements with those of sea level in order to capture daily and sub-  
243 daily tides at Cape-Verdian latitudes. For this purpose, we used hourly data from the tide gauge  
244 station Palmeira (16.75°N, -22.983333°E), located on Sal island (Cape Verde, Figure 1b). The data  
245 are acquired and validated by the UH Sea Level Center via the Permanent Service for mean sea  
246 level platform (Holgate et al., 2013; PSMSL, 2019).

247

### 248 *3.4 Singular Spectrum Analysis*

249 We analyzed the five geophysical time-series using the Singular Spectrum Analysis (SSA). SSA is a  
250 non-parametric time-series analysis technique based on two steps, a decomposition and a  
251 reconstruction allowing to identify trends, periodic or quasi-periodic components and noise in any  
252 signal as independent and physical components (Vautard and Ghil, 1992; Golyandina and  
253 Zhigljavsky, 2013). The first step consists in building a diagonal auto-correlation matrix  
254 corresponding in our case to a Hankel matrix. To do so, lagged copies representing segments of  
255 length L of the time-series are embedded to create a L x K matrix, with K = N -L +1 and N the  
256 length of the time-series. This analysis is sensitive to the dimensions of the Hankel matrix as  
257 smaller L will make the analysis closer to fractal, while larger L will result in reconstructions very  
258 similar to the original time-series. The Hankel matrix, H, is then decomposed using the Singular  
259 Value Decomposition (SVD) algorithm such as follows:

$$H = \sum_{i=1} \sqrt{\lambda_i} U_i V_i^T$$

261 Where  $\sqrt{\lambda_i}$  is the matrix containing the singular values,  $U_i$  and  $V_i$  the left and right matrices of the  
262 orthogonal singular vectors. One of the advantages of the SVD is that it ranks the singular values in  
263 the descending order of the energies and therefore their contribution to the original time-series as  
264 well. This way, a signal whose mean value is non-null is characterized by a trend that will appear in  
265 the first eigenvalues. The eigen triplets (eigenvalues and eigenvectors) associated with eigenvalues  
266 of similar magnitude may be grouped. In the case of a perfect oscillation, SSA produces two  
267 identical eigenvalues called Hilbert pair. Thus, different components of the time-series, e.g the  
268 trend, (pseudo) cycles and noise can be identified and reconstructed by gathering small groups of  
269 eigenvalues having very close contributions. Noise is usually of small amplitude and represented by  
270 small eigenvalues. Identifying the eigenvalues of interest allows to filter non-linearly the original  
271 signal by truncating the SVD using the selected eigenvalues.

272 In the present study, we focus on components whose periods were very similar to that of known  
273 Earth tides and whose separability with neighboring components was sufficient to properly identify  
274 individual tidal components. The uncertainties on each identified period as later presented in  
275 [Figures 3 & 4](#) and [Table 1](#), were estimated using the half-width of their peaks at half-height,  
276 similarly to [Le Mouél et al. \(2019\)a,b](#) and [Dumont et al. \(2020\)](#). More details on the method, its  
277 application and discussion on critical aspects of SSA can be found in [Lopes et al. \(2017\)](#) and [Le](#)  
278 [Mouél et al., \(2019a,b; 2020\)](#).

279 The significant cloud cover associated with lower lava emissions from mid-December to  
280 February 2015 originated important gaps in the estimation of lava VFR and SO<sub>2</sub> VCD measured  
281 from SEVIRI data. SSA was therefore applied from 23th November to 22 December 2014, covering  
282 a 30-days interval, for the lava VFR and SO<sub>2</sub> time-series. For similar reasons, we analyzed the  
283 seismic tremor detected at FCVT station for the period starting on 30<sup>th</sup> November and ending on 26  
284 December 2014 ([Figure 2](#)). The l.o.d and sea level time-series representing reference data sets for  
285 identifying tidal periodicities, were analyzed using SSA for the whole eruption duration (23  
286 November 2014 – 8 February 2015).

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### 290 3.5 Tidal potential calculation

291 The tidal potential  $W$  (see [Jobert and Coulomb, 1973, chapter 18](#)) is given by the following

292 expression :

$$W = \frac{GM'}{D} \sum_{n=2}^{\infty} \left(\frac{R}{D}\right)^n P_n(\cos(z))$$

293 constant,  $M'$  the mass of the celestial body, e.g that of the Sun or the Moon;  $D$  its distance to the

294 Earth;  $R$ , the Earth's radius and  $P_n$  the Legendre's polynomial function of the cosine relating to the

295 celestial body position.  $W$  may be decomposed in a sum of spherical harmonics. It is usually

296 reduced to its first term ( $n=2$ ) when focusing on the strongest tides induced by the gravitational

297 effect of Moon and the Sun, the distances between the other planets of the solar system and the

298 Earth being tremendous.  $W_2$  may be described as follows ([Jobert and Coulomb, 1973](#)):

$$299 W_2 = (GM'R^2)(P_2 \cos Z) / D^3$$

300 with  $Z$ , the zenithal distance of the celestial body at the observation site on Earth and  $P_2$  the

301 Legendre's polynomial of second order (see graphic [Figure 5a](#)).  $D$  and  $Z$  vary in time for both the

302 Sun and the Moon.  $D$  time-series are freely accessible from the Institut de Mécanique Céleste et de

303 calcul des Éphémérides website (<http://vo.imcce.fr/webservices/miriade/?forms>). The zenithal

304 distance  $Z$ , may be calculated using the right ascension ( $H$ ) and the declination ( $\delta$ ) of the Moon/Sun

305 that is given by the ephemeris, as well as the longitude ( $\Phi$ ) and latitude ( $\Theta$ ) of the observation point.

306 It follows that, at about one common factor given by  $GM'R^2/D^3$ , the tidal potential  $W_2$  is composed

307 of three main tides, the zonal ( $W_{2,0}$ ), tesseral ( $W_{2,1}$ ) and sectorial ( $W_{2,2}$ ) components which are

308 characterized by different periods and specific spatial distribution. They can be expressed as follows

309 ([Jobert and Coulomb, 1973](#)) :

$$\begin{aligned}
W_{20}(\cos Z) &= 0.5(3\cos^2\theta - 1) * 0.5(3\cos^2\delta - 1) \\
W_{21}(\cos Z) &= 3\sin\theta\cos\theta * 3\sin\delta\cos\delta * \cos(H - \phi) \\
W_{22}(\cos Z) &= 3\cos^2\theta * 3\cos^2\delta * \cos 2(H - \phi)
\end{aligned}$$

311 With these equations, it is clear that the spatial configuration of the Moon and Sun with respect to  
312 the observation site at a specific time will define the magnitude and phase shift of the tidal potential  
313 experienced at the observation site.

314

#### 315 4. Results

316 MSG-SEVIRI satellite images reveal a maximum of lava VFR and SO<sub>2</sub> VCD detected  
317 between 1 and 2 days respectively, after the onset of the eruption (Figure 2). Although the seismic  
318 tremor data does not cover the first week of the eruption, it still shows significant variations with a  
319 maximum reached on 3<sup>rd</sup> December, followed by an overall slow decay. A stronger decline is  
320 observed for the SO<sub>2</sub> VCD that drops by half in about a week and then slightly varies around an  
321 index value of 8 for the rest of the eruption. The lava VFR shows a different behavior with a strong  
322 emission of lava including large variations up to mid-December, reported as cycles by Cappello et  
323 al., (2016). For this time interval, the l.o.d. shows regular variations of the order of 0.5 ms (Figure  
324 2), except in the first week of January 2015 where a drop exceeds 0.8 ms. The sea level time-series  
325 show larger variations of the order of 140-160 cm on average, with an overall trend similar to that  
326 of the l.o.d (Figure 2). Both l.o.d. and sea level variations are governed by clear periodic oscillations  
327 that span different time scales as illustrated by their spectra (Figure 3a). Normalized spectra already  
328 reveal common periodicities in the three volcanological time-series of the first month of the  
329 eruption what we will further explore next.

330

331 We analyzed the trends and (pseudo-) periods that compose the lava VFR, SO<sub>2</sub> and seismic  
332 tremor time-series by considering up to 20-25 eigenvalues for each SSA decomposition (Figure 3b).  
333 We thus identified 9 periods corresponding to the main components of the Earth tides, ranging from  
334 0.5 to ~15 days and an additional ~4.5-day period (see Table 1). The presence of this ~4.5-day

335 period oscillation in two of the three geophysical data series, e.g SO<sub>2</sub>, lava VFR and l.o.d. likely  
336 reflects a common origin, what we interpret as a harmonic of a longer lunisolar or lunar tide. Thus,  
337 we clearly detected 5 tidal components in the seismic tremor, 5 in the SO<sub>2</sub> time-series and 4 in the  
338 lava VFR, which match well those identified in the l.o.d. or sea level data (Table 1). One can note  
339 the clear dominance of lunar tidal periods over those of solar origin. For each volcanological time-  
340 series, we combined these tidal components with the trends that are usually associated with volcanic  
341 processes (see Dumont et al., 2020) to reconstruct the original time-series (Figure 3c). We only  
342 excluded the solar semi-diurnal component (S2) extracted from the lava VFR data, as it is not a pure  
343 tide but rather a tide-induced thermal effect on the atmosphere. Indeed, the absorption of solar  
344 radiations produces water vapor that circulates and condensates in the atmosphere, releasing thus  
345 latent heat during the water vapor cycle at harmonics of a solar day (Hagan and Forbes, 2002).  
346 These trends and lunisolar tidal periods that we identified allow to reproduce most variations as  
347 illustrated with the Figure 3c. More precisely, we are able to reconstruct up to 96.2% of the SO<sub>2</sub>  
348 time-series, 85.6 % of the seismic tremor and 53.1% of the lava VFR. Next, we present some of the  
349 waveforms associated with the tidal components identified in the volcanological time-series.

350         Figure 4 presents two of the three long tides extracted in the seismic tremor, lava VFR and  
351 SO<sub>2</sub> time-series. The seismic tremor displays two out of these three long-period tides (Table 1). The  
352 component with a period of 14.9 +/- 5.8 days, likely corresponds to the lunisolar fortnightly tide,  
353 being possibly caused either by the oscillations of the Moon declination, e.g Mf with a 13.6-days  
354 period, or reflecting half of the synodic lunar month, e.g MSf with a 14.7-days period. Despite a  
355 relatively large uncertainty, this tidal component identified in the seismic tremor appears clearly in  
356 phase opposition with that of the l.o.d (Figure 4). The lunar constituent Mt with a ~7.0-days period,  
357 is in-phase for both seismic tremor and lava VFR. For both components, the amplitude increases up  
358 to ~10th December, after which it decreases, a pattern that is more pronounced for the seismic  
359 tremor. Figure 4 also shows shorter periods with the example of one diurnal and one semi-diurnal.  
360 Four tidal constituents with near diurnal periods are detected in the volcanological time-series,

361 namely 2Q1, M1, P1 and K1 (Table 1). The P1 tidal component detected in the lava VFR may  
362 possibly reflect the sensitivity of the mid-wave infrared band (3.9  $\mu\text{m}$ ) to solar diurnal tides, even  
363 though data processing includes a step to remove it. Except for the solar component P1 detected  
364 only in the lava VFR, the three other near-diurnal constituents were detected in both the SO<sub>2</sub> and  
365 seismic tremor as well as in the sea level data. The two components of the larger elliptic diurnal or  
366 2Q1 tide are mainly in phase opposition and show a strong attenuation starting early December for  
367 the SO<sub>2</sub> and after 21<sup>st</sup> December for the seismic tremor (Figure 4). Although centered on 1.17 days,  
368 the SO<sub>2</sub> component of the 2Q1 tide has an uncertainty that is wide enough to allow it to be  
369 interpreted as Q1 or M1. The lunar elliptic semi-diurnal components (N<sub>2</sub>, first or second order,  
370 Table 1) in the SO<sub>2</sub> and seismic tremor time-series also attenuate over time. They are decaying at a  
371 very similar rates, although they are in quadrature after 1<sup>st</sup> December (Figure 4). One can note that  
372 the amplitude of the SO<sub>2</sub> and seismic tremor components are larger when those of the sea-level are  
373 smaller.

374

## 375 5. Discussion

376 The Singular Spectrum Analysis applied to three geophysical time-series acquired during the  
377 2.5-month Fogo eruption provides a new evidence of the influence of lunisolar gravitational forces  
378 on effusive eruptions. We identified between 4 and 5 tidal periodicities in the SO<sub>2</sub>, lava VFR and  
379 seismic tremor time-series with periods ranging from semi-diurnal to fortnightly (Table 1, Figure 4).  
380 These results confirm and complement the observations made by Dumont et al. (2020) showing that  
381 the movements of magmatic fluids at volcanoes (melt, erupted magma and gas emission) are not  
382 modulated by one but rather by a combination of several earth tides as it had also been suggested by  
383 Sottili and Palladino (2012) and Bredemeyer and Hansteen (2014). In addition, this analysis  
384 confirms the predominant influence of lunar tides over lunisolar ones on volcanic processes, e.g De  
385 Lauro et al., (2013).

386 It is interesting to note that this analysis detected many more tides of short periods ([Table 1](#))  
387 than those that had been detected for Holuhraun eruption that was ongoing before the onset and  
388 after the termination of Fogo eruption ([Dumont et al., 2020](#)). This observation might however be  
389 biased by the time interval considered by this study, i.e.  $\sim 1$  month instead of the  $\sim 2.5$  months of the  
390 whole eruptive period at Fogo, preventing the detection of longer tides. Another way to address the  
391 differences of tides present in the co-eruptive geophysical time-series in Iceland and Cape Verde is  
392 to consider the components identified in each time-series and the rank of their associated  
393 eigenvalues ([Figure 3b](#)). The longest tide we identified for Fogo eruption, the fortnightly, appears in  
394 the seismic tremor only as fourth component, after the trend, a 2.4-day period and a 7.9-days period  
395 oscillations. This contribution of Mf/MSf to the seismic tremor time-series at Fogo is smaller than  
396 those in both the seismic tremor and radiated power for the Holuhraun eruption (see the values of  
397 the normalized eigenvalues,  $\sim 0.05$  versus  $\sim 0.07$ , [Figures 3](#) this study and [Dumont et al., 2020](#)), that  
398 are also systematically ranked second after the trend. Similarly, the 9.4 day-period or Mtm tide, was  
399 only detected in the SO<sub>2</sub> data at Fogo, where it appears in tenth position ([Figure 3b](#)). This tide was  
400 ranked higher in the seismic tremor of the Holuhraun eruption, in sixth position, associated with  
401 higher normalized eigenvalues ([Dumont et al., 2020](#)). The semi-diurnal and diurnal oscillations  
402 detected in the three geophysical time-series during Fogo eruption are more numerous, with ranks  
403 between 4 and 25 ([Figure 3b](#)). This comparison suggests that at low latitude, longer tides are  
404 generally weaker than at higher latitudes and have also a stronger contribution in parameters related  
405 to magma propagation, as shown by VFR and seismic tremor time-series, rather than those related  
406 to gas fluxes. This variety of lunar, solar and lunisolar tides was not detected for the Icelandic data  
407 sets. It is important to remind that the tidal contributions we captured for Fogo eruption are of first-  
408 order as we could only analyze slightly less than half of the eruption duration.

409 Let us now further explore the contribution of the fortnightly component at these two  
410 eruption sites, Fogo and Holuhraun, by calculating the tidal potential. The fortnightly periodicity Mf  
411 is related to variations of the Moon declination which are mainly related to the zonal component of

412 the tidal potential,  $W_{2,0}$ . We estimated  $W_{2,0}$  between late November and late December 2014 at the  
413 Icelandic and Cape-Verdian eruption sites (Figure 5). This calculation clearly shows that this tide is  
414 stronger at higher latitude than nearby the Equator explaining why it was more easily and  
415 systematically detected in all geophysical parameters for the Holuhraun eruption than for Fogo  
416 (Dumont et al., 2020). Moreover, because of its location in the Equatorial band, Fogo volcano is  
417 more sensitive to the solar tides than Iceland, and even more so towards the end of the year when  
418 the Earth is closest to the Sun. Altogether, these observations suggest that the influence that tides  
419 have on volcanic activity is a result of the combination of tides that is specific to the volcano  
420 location on Earth and also to the timing of the eruption which is associated with a specific  
421 astronomical configuration. We also note that the onset of Fogo eruption dated to the 23<sup>rd</sup> November  
422 2014 occurred only one day after a syzygy (new Moon). Such a short timing have been evoked to  
423 suggest the tidal influence on the triggering mechanisms of volcanic eruptions (Mauk and Johnston,  
424 1973; Dzurisin, 1980; Jentzsch et al. 2001; Sottili and Palladino, 2012; Dumont et al., 2020).

425 The three time-series that we analyzed for Fogo eruption provide insights on 1) the  
426 movement of volatile-bearing magma from depth to surface, as documented by the seismic tremor,  
427 2) the emission of gas once the magma erupted, as evidenced with the SO<sub>2</sub> emissions and 3) the  
428 movement of magma at Earth's surface as illustrated with the lava VFR. Although tidal waves do  
429 not require a medium to propagate as they deform the space-time geometry, the waveforms  
430 associated with the tidal components that we extracted from the volcanological time-series seem to  
431 show some features related to their propagation medium. Actually, attenuation of the waveforms is  
432 systematically observed for the seismic tremor and the SO<sub>2</sub> emission in particular, but (almost) not  
433 observed in the lava VFR (Figure 4). Indeed, the lava field growth was characterized at Fogo by  
434 early formation of lava tubes that started early December for the second and main flow and earlier  
435 for the first and short flow (Calvari et al., 2018). The efficient thermal insulation of lava tubes  
436 (Keszthelyi, 1995) as well as the steady magma supply rate (Calvari et al., 2018) could explain this  
437 reduced attenuation of lava VFR waveforms. The seismic tremor is thought to be associated with

438 movement of fluids at subsurface and has often be correlated with the intensity of volcanic activity  
439 (Chouet, 1985; Eibl et al., 2017 and therein references). The attenuation observed for the tidal  
440 components of the seismic tremor may represent the decrease of the eruption vigor. However, the  
441 ascent of a hot and multi-phase melt in a very heterogeneous medium (fractured and temperature-  
442 eroded) towards the surface at rates that are likely varying over the eruption duration, may also  
443 contribute to attenuate differently the various tidal components of the seismic tremor. Finally, the  
444 strong attenuation detected in the gas emissions is likely due to the dispersive nature of the  
445 atmosphere.

446

447 All these observations suggest that volcano's response to Earth's tides is highly dependent on the  
448 nature of the magma erupted, as well as on the structure of the internal plumbing system and on  
449 tectonic setting, as also previously suggested by various studies (Mauk and Jonhston, 1973;  
450 Dzurisin, 1980; Sparks, 1981; Petrosino et al., 2020). To further characterize the volcano's response  
451 to this quasi-permanent external forcing, let's consider  $x$  as the tidal potential and  $y$  as the  
452 geophysical parameters influenced by earth tides and recorded during an eruption (e.g. volcanic  
453 tremor, gas release, lava emission...) and  $h$ , the impulse response of the volcano corresponding to all  
454 geochemical-physical processes characterizing the volcano activity. Using filter theory, we can  
455 describe the impulse response of the volcano as follows:

456

$$y(t) = (h * x)(t)$$

457 In the Fourier space, this equation corresponds to a simple product:

$$Y(f) = H(f) \cdot X(f)$$

458  $Y$  and  $X$  are the Fourier transforms of  $y$  and  $x$  respectively and  $H$  corresponds to the transfer  
459 function of the volcano, with  $f$ , the frequency. SSA allowed us to extract and identify the pseudo-  
460 oscillations of tidal origin in the three geophysical time-series. It is therefore possible to reconstruct  
461  $H$  for all the identified periodicities. The tidal potential  $x$  for the identified tidal periods is known  
462 (Guinot, 1973; Ray et al., 2014). By representing the amplitude and phase of each tidal period  
463 extracted from the seismic tremor, lava VFR and  $\text{SO}_2$  as a function of their frequencies, we can

464 characterize the pulse response of Fogo volcano. [Figure 6](#) shows that Fogo volcano seems to  
465 respond as a bandpass filter ( $\sim 0.1\text{-}0.85\text{ days}^{-1}$ ) to the external tidal forcing, being more sensitive to  
466 tidal periods slightly longer than 24 h such as those of Q1, and up to longer tides with periods up to  
467  $\sim 9$  days ([Table 1](#)), including Msp, Mt, Mtm tides. We suggest that the band width of this filter is  
468 likely related to the internal structure of the volcano, its chemistry and geodynamical setting, as  
469 suggested by [Mauk and Jonhston \(1973\)](#), [Dzurisin \(1980\)](#) or [De Lauro et al., \(2018\)](#). Further  
470 investigations on such response of volcanoes to external forcing could provide a new way to  
471 address the interplay between internal and external processes that influence magmatic and volcanic  
472 activity.

473

## 474 **6. Conclusion**

475 Our study focuses on the 2014-2015 eruption of the Fogo volcano, Cape Verde. We analyze three  
476 co-eruptive geophysical time-series, namely the seismic tremor, SO<sub>2</sub> emissions and lava volume  
477 flow rate (VFR), using the Singular Spectrum Analysis (SSA). By considering the first month of the  
478 eruptive activity, we were able to identify between 4 and 5 different tidal periods in each of these  
479 volcanological time-series, ranging from semi-diurnal to fortnightly periods. These results clearly  
480 show a predominant influence of lunar tides, although solar and lunisolar tidal periods were also  
481 detected. The retrieval of the waveforms associated with the tidal components extracted from the  
482 different time-series, reveals a stronger attenuation for the tidal components of the SO<sub>2</sub> and seismic  
483 tremor. By comparing these observations to those obtained for Holuhraun eruption (Iceland)  
484 ([Dumont et al., 2020](#)), we illustrate that the forcing of tides upon volcanic activity is a result of both  
485 the volcano location on Earth and of the timing of the eruption, which corresponds to a specific  
486 astronomical configuration. Finally, using filter theory we demonstrate that Fogo volcano responds  
487 as a bandpass filter to the quasi-permanent tidal oscillations.

488

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500

#### 501 **Authors' contributions**

502 SD: Formal analysis; Visualization; Writing - original draft;

503 GS, SC and MG: Formal analysis; reviewing and editing;

504 FL and JLLM: Conceptualization; Methodology; editing;

505 YG: Formal analysis;

506

507

#### 508 **Data Availability**

509 The SO<sub>2</sub> emission and lava volume flow rate (VFR) time-series are available from the HOTVOLC  
510 platform (<https://hotvolc.opgc.fr>). The l.o.d time-series is freely accessible as part of the EOP14C04  
511 data set provided by the International Earth Rotation Service ([https://www.iers.org/iers/en/DataProducts/](https://www.iers.org/iers/en/DataProducts/EarthOrientationData/eop.html)  
512 [EarthOrientationData/eop.html](https://www.iers.org/iers/en/DataProducts/EarthOrientationData/eop.html), IERS, Paris, France) as well as that of the sea level, accessible from  
513 Permanent Service for mean sea level platform (PSML, 2019). The seismic tremor data is available  
514 upon request.

515

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721 **DATA REFERENCES**

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**FIGURES**

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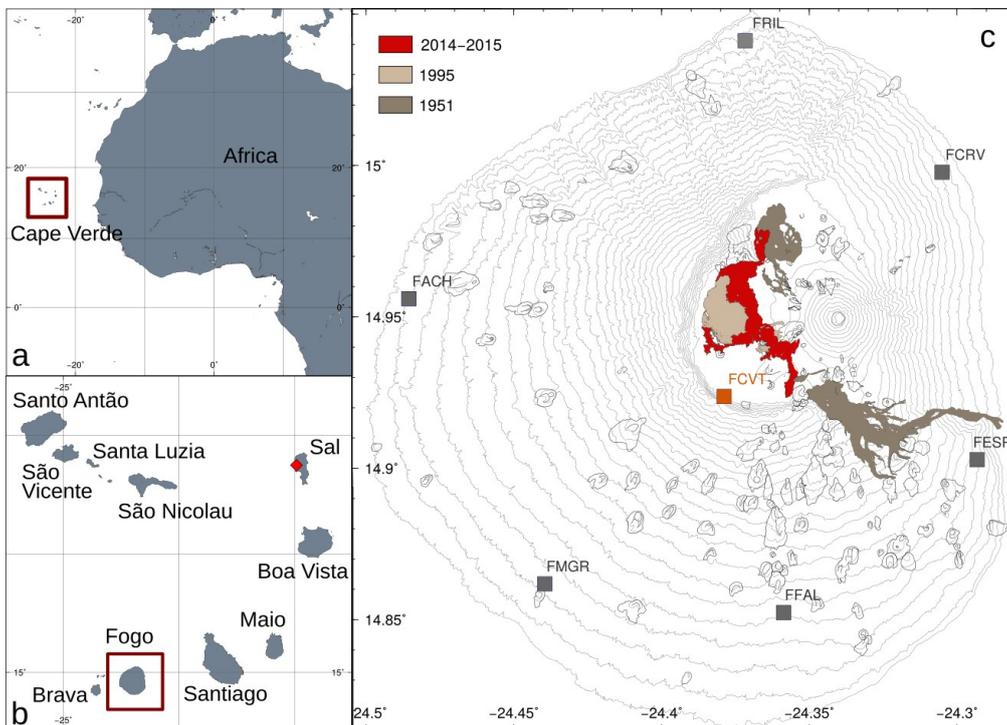
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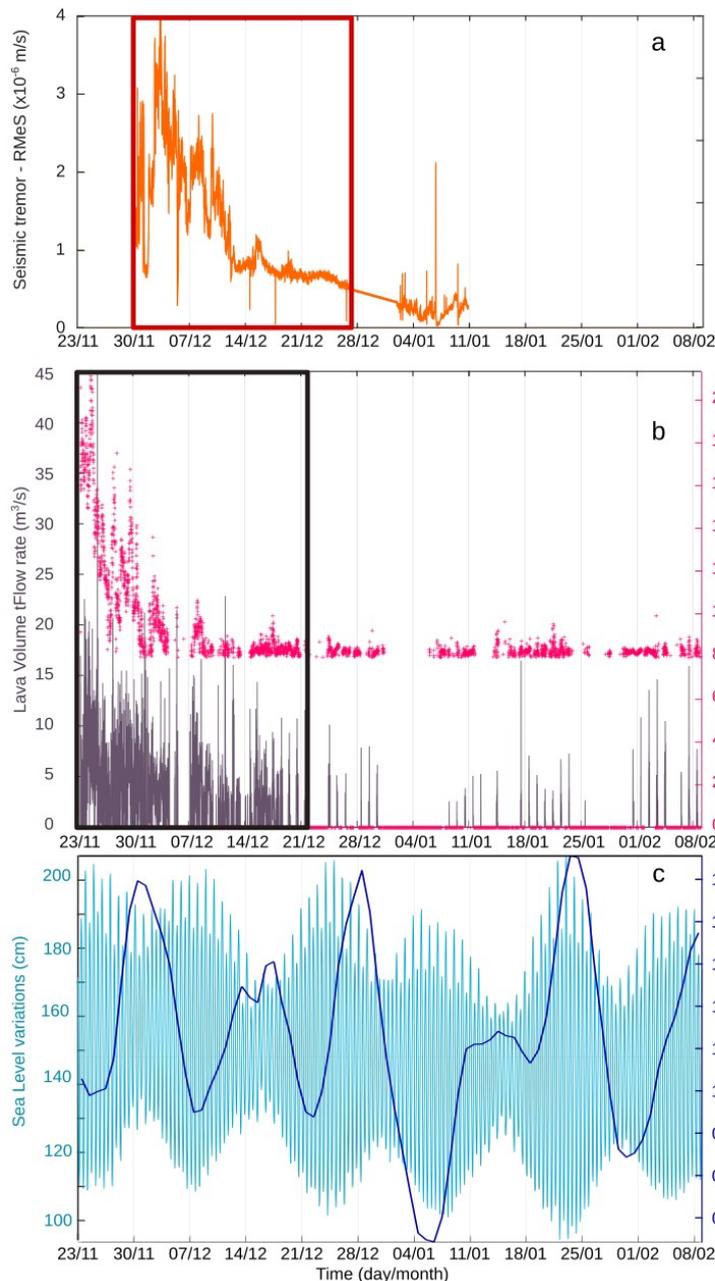


744 **Figure 1:** Map of Fogo volcano and the Cape Verde archipelago. Located ~500 km west of the

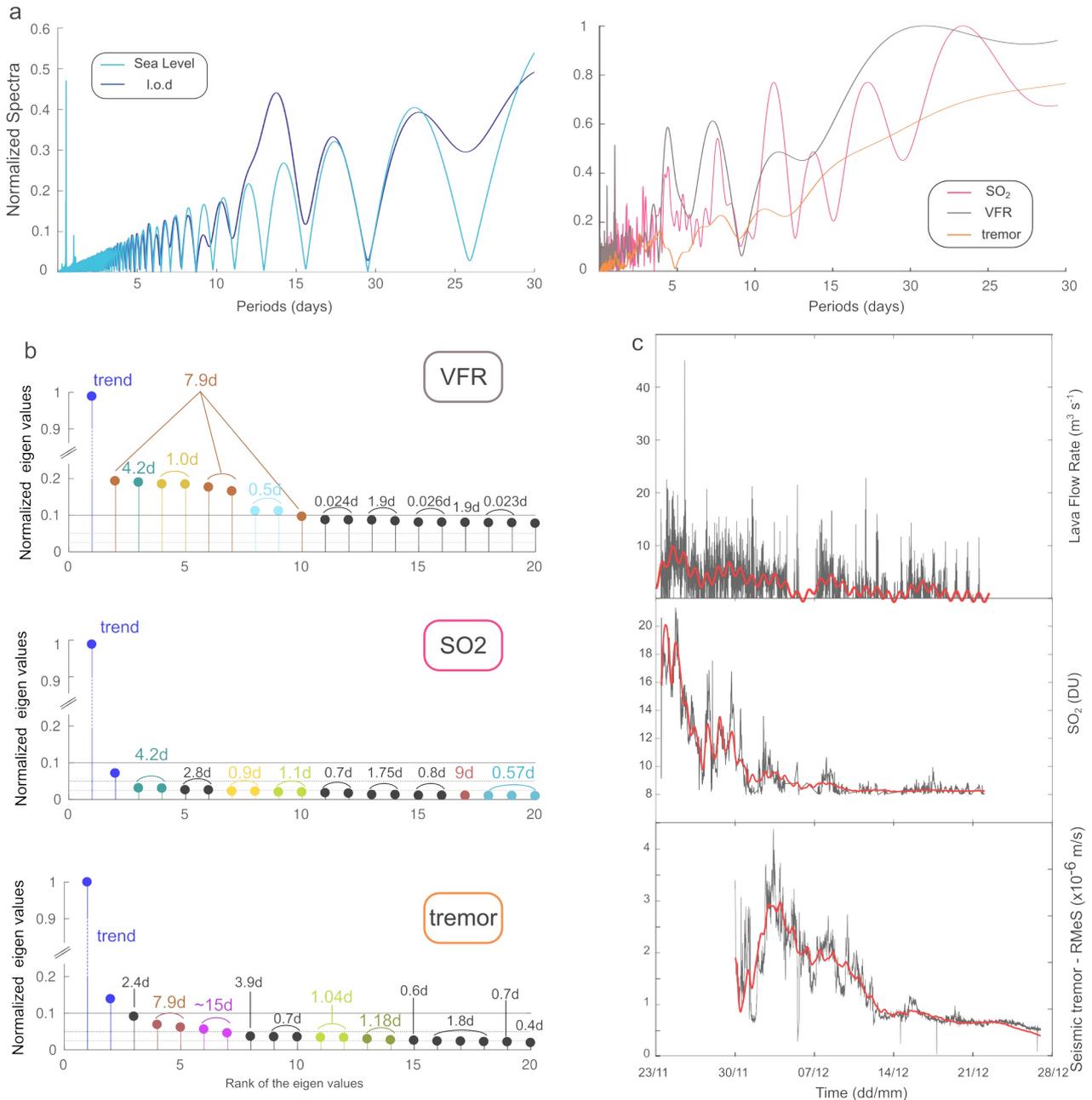
745 African coast (a), Cape Verde archipelago is composed of a dozen of islands forming a westward-

746 open crescent (b). Fogo island is located on the southern branch. The red diamond represents the  
747 tide gauge where sea level measurements were acquired and used for this study. Fogo island is  
748 characterized by a central caldera, Chã das Caldeiras, open to the east (c). Fogo volcano rises  
749 nearby the caldera center. The lava fields produced during the last century are represented by the  
750 color codes indicated on the upper left of the figure (Torres et al., 1998; Vieira et al., 2020). The  
751 topography is represented using 100-m isopachs. The seismic station FCVT (orange square) used in  
752 this study was part of a seismic network (gray squares) deployed by the Portuguese government and  
753 the C4G consortium during the 2014-2015 eruption.

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773 **Figure 2:** Time-series of the seismic tremor, SO<sub>2</sub> emissions and lava VFR recorded during the  
 774 2014-2015 Fogo eruption. (a) Seismic tremor was detected in the 1-4 Hz band at the FCVT station  
 775 (see Figure 1c for location). (b) SO<sub>2</sub> emission and lava VFR derived from the MSG-SEVIRI  
 776 satellite data and (c) length-of-day and sea level time-series spanning the whole eruption. The two  
 777 rectangles indicate the time intervals considered for the present study.

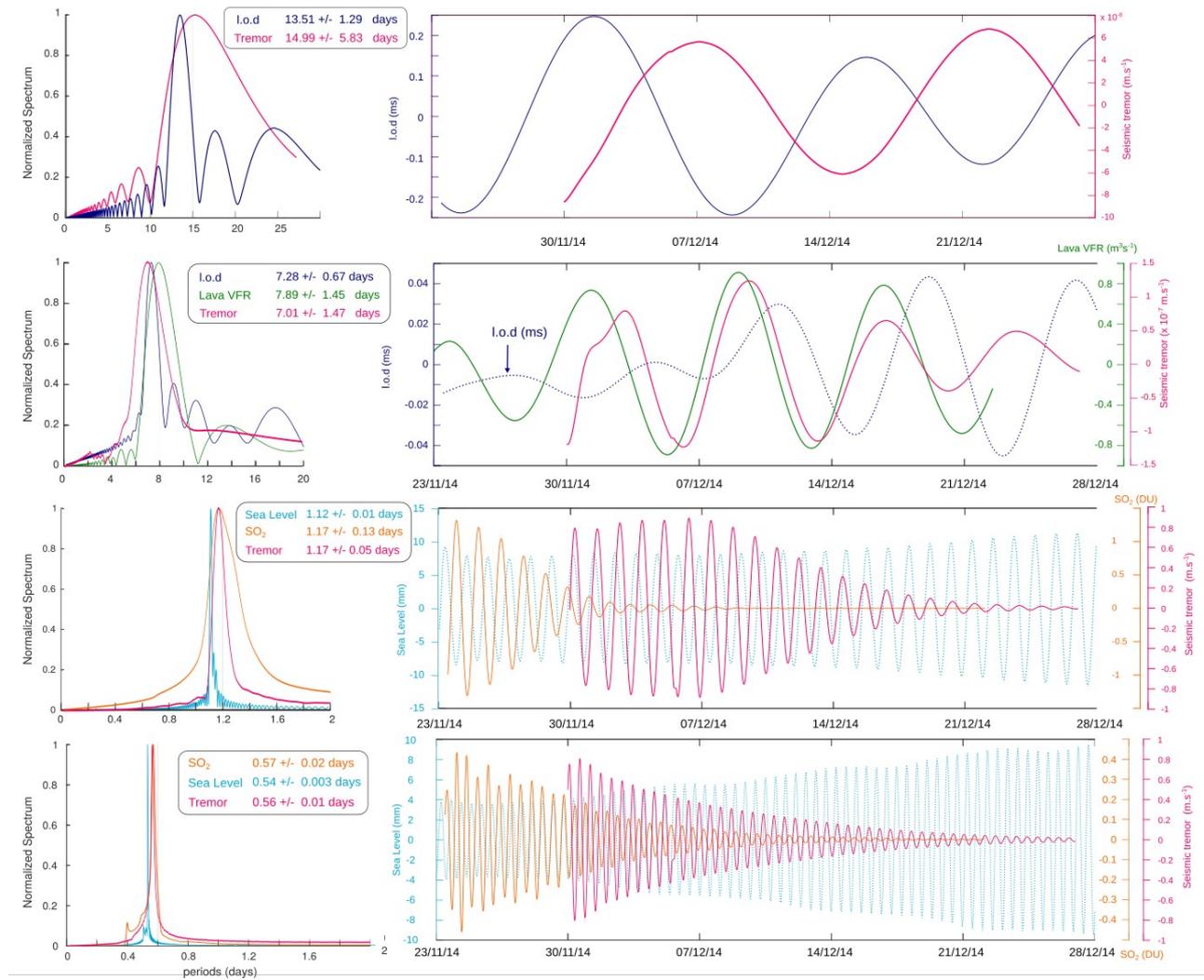


778 **Figure 3:** Spectral content and reconstruction of the original volcanological time-series. (a)  
 779 Normalized spectra of the l.o.d and sea level time-series (left) and for the seismic tremor, SO<sub>2</sub> and

780 lava VFR time-series (right). (b) Eigen values ranked in decreasing order, are gathered as  
 781 components representing the trends and (quasi) periodic oscillations of each time volcanological  
 782 time-series. The different colors correspond to different periodicities identified in the lava VFR, SO<sub>2</sub>  
 783 and seismic tremor time-series. The black ones were not recognized as main tides and not further  
 784 considered in this study. (c) Reconstruction (red) of the original time-series (black) of the lava VFR  
 785 (up), SO<sub>2</sub> (middle) and seismic tremor (bottom) using the trends and the different tidal (pseudo)-  
 786 periods identified using SSA shown in (b) and listed in Table 1.

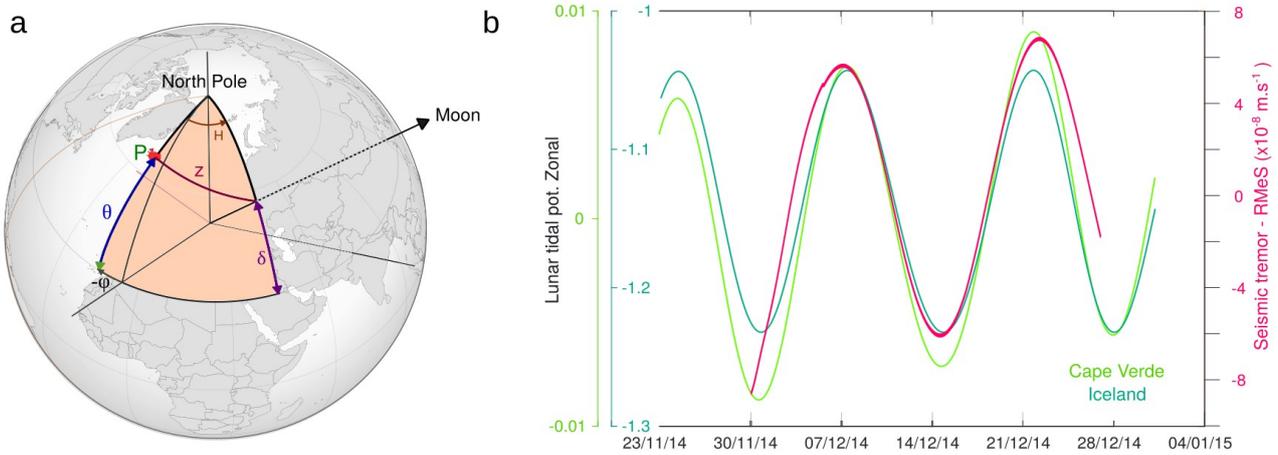
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790 **Figure 4:** Examples of tidal components extracted in the seismic tremor, SO<sub>2</sub> and lava VFR time-  
 791 series from ~15 to ~0.5 days (from top to bottom). The normalized spectra are shown on the left  
 792 side and the corresponding waveforms on the right one. The tidal components of the seismic tremor  
 793 are represented in pink, those of the SO<sub>2</sub> emission in orange and the lava VFR in green. The  
 794 components extracted from the l.o.d are in dark blue and in light blue for the sea level, sometimes  
 795 represented using dotted lines for clarity.

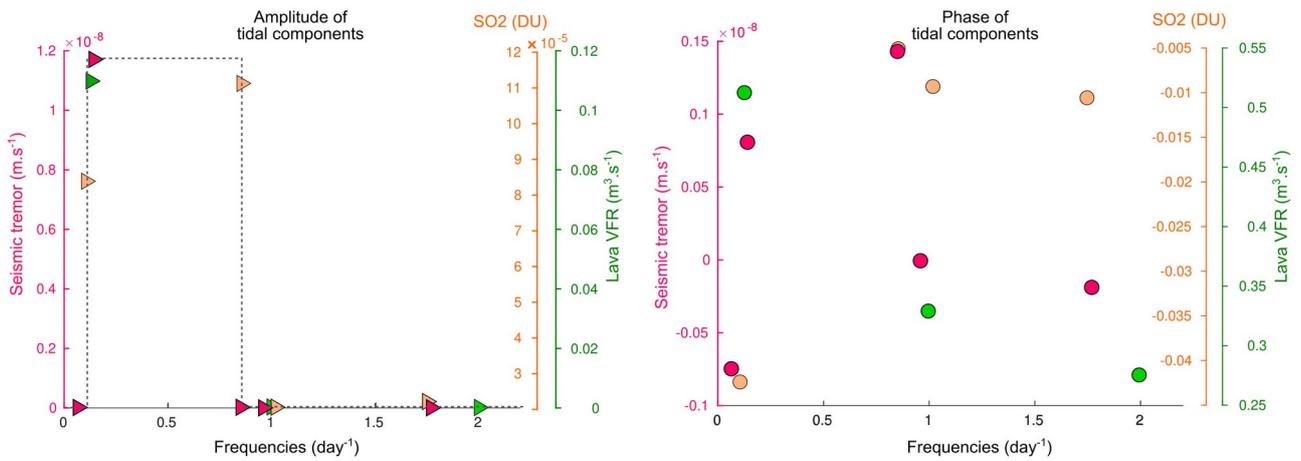
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 801 **Figure 5:** Zonal component of the lunar tidal potential. (a) Parameters used when calculating the  
 802 tidal potential, depending on the position of the observation site on Earth's planet and that of the  
 803 celestial body inducing the gravitational force, as the Moon for instance. See Section 3.5 for details.  
 804 (b) The zonal component due to the Moon gravitational attraction was estimated for two  
 805 observation sites, Iceland ( $\theta = 65^\circ$ ) and Cape Verde ( $\theta = 15^\circ$ ). These zonal components are  
 806 compared to the fortnightly component extracted in the seismic tremor at Fogo.

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818 **Figure 6:** Amplitude and phase associated with all tidal components extracted in the three  
819 volcanological time-series during the Fogo eruption, namely seismic tremor (pink), SO<sub>2</sub> emission  
820 (orange) and lava VFR (green). The volcano acts as a bandpass filter to tidal components showing  
821 sensitivity in the frequency band  $\sim 0.1$  to  $0.85 \text{ days}^{-1}$ , and filtering out all other tidal components, as  
822 shown by the dotted brown line.

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I.o.d (days)	sea level (days)	seismic tremor (days)	lava VFR (days)	SO <sub>2</sub> (days)	Earth tide names	Origin L: Moon; S: Sun
13.51 ± 1.29		14.99 ± 5.8 ( <i>0.07</i> )			Mf or MSf	L or L+S
9.01 ± 0.57				9.40 ± 2.30 ( <i>0.11</i> )	Mtm	L
7.28 ± 0.67		7.01 ± 1.5 ( <i>0.14</i> )	7.89 ± 1.45 ( <i>0.14</i> )		Mt	L
4.05 ± 0.15*			4.28 ± 0.42*	4.46 ± 0.97*		
	1.12 ± 0.01	1.17 ± 0.05 ( <i>0.85</i> )		1.17 ± 0.13 ( <i>0.86</i> )	2Q <sub>1</sub>	L
	1.08 ± 0.009	1.03 ± 0.1 ( <i>0.96</i> )			M <sub>1</sub>	L
			1.002 ± 0.02 ( <i>1.0</i> )		P <sub>1</sub>	S
	0.999 ± 0.008			0.980 ± 0.06 ( <i>1.02</i> )	K <sub>1</sub>	L + S
	0.54 ± 0.003	0.56 ± 0.01 ( <i>1.78</i> )		0.57 ± 0.02 ( <i>1.75</i> )	2"N <sub>2</sub> or N <sub>2</sub>	L
			0.50 ± 0.005 ( <i>2.0</i> )		S <sub>2</sub>	S

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830 **Table 1:** Tidal periodicities identified in the three volcanological time-series, e.g. seismic tremor,  
831 lava VFR and SO<sub>2</sub>, also extracted in the length-of-day and sea-level variations, used as proxy of the  
832 tidal action on Earth. The numbers indicated into brackets in italic are the corresponding  
833 frequencies to tidal periodicities in day<sup>-1</sup>, as represented in Figure 6.